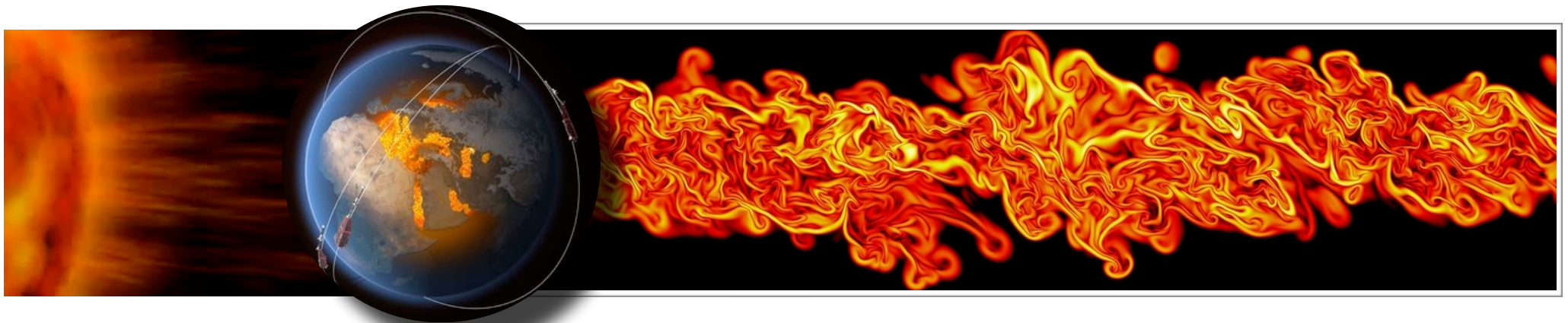




A Study of Ionospheric Turbulence in the Polar Regions by Swarm constellation



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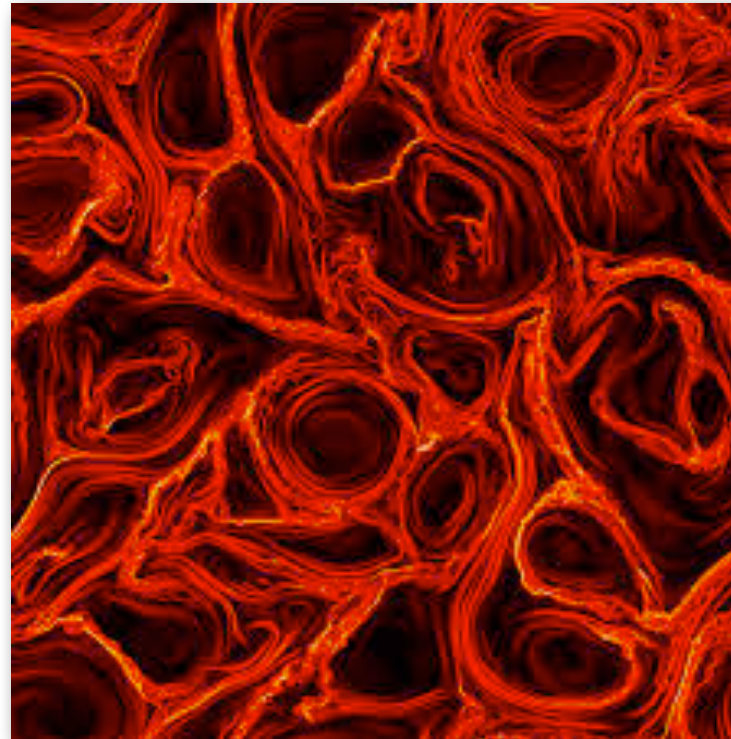


AIM

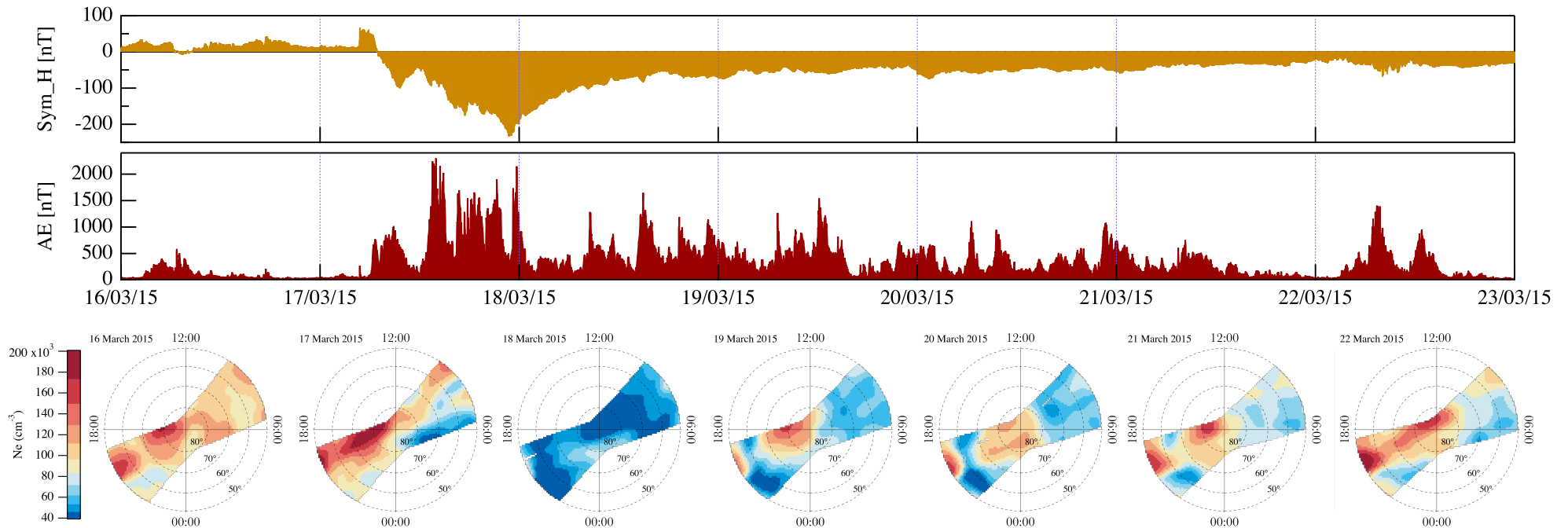
★To analyse the **scaling features of the electron density** fluctuations in order to get information about **the role of turbulence in the ionospheric** during a geomagnetic storm.

WHY

Turbulence plays a central role in several processes involving the plasma and particle transport. In the Earth's ionosphere, it plays a fundamental role in the overall dynamics. It is able to generate/create magnetic and plasma structures that can strongly affect the homogeneity of the plasma in the ionospheric regions. It can play a relevant role in space weather.

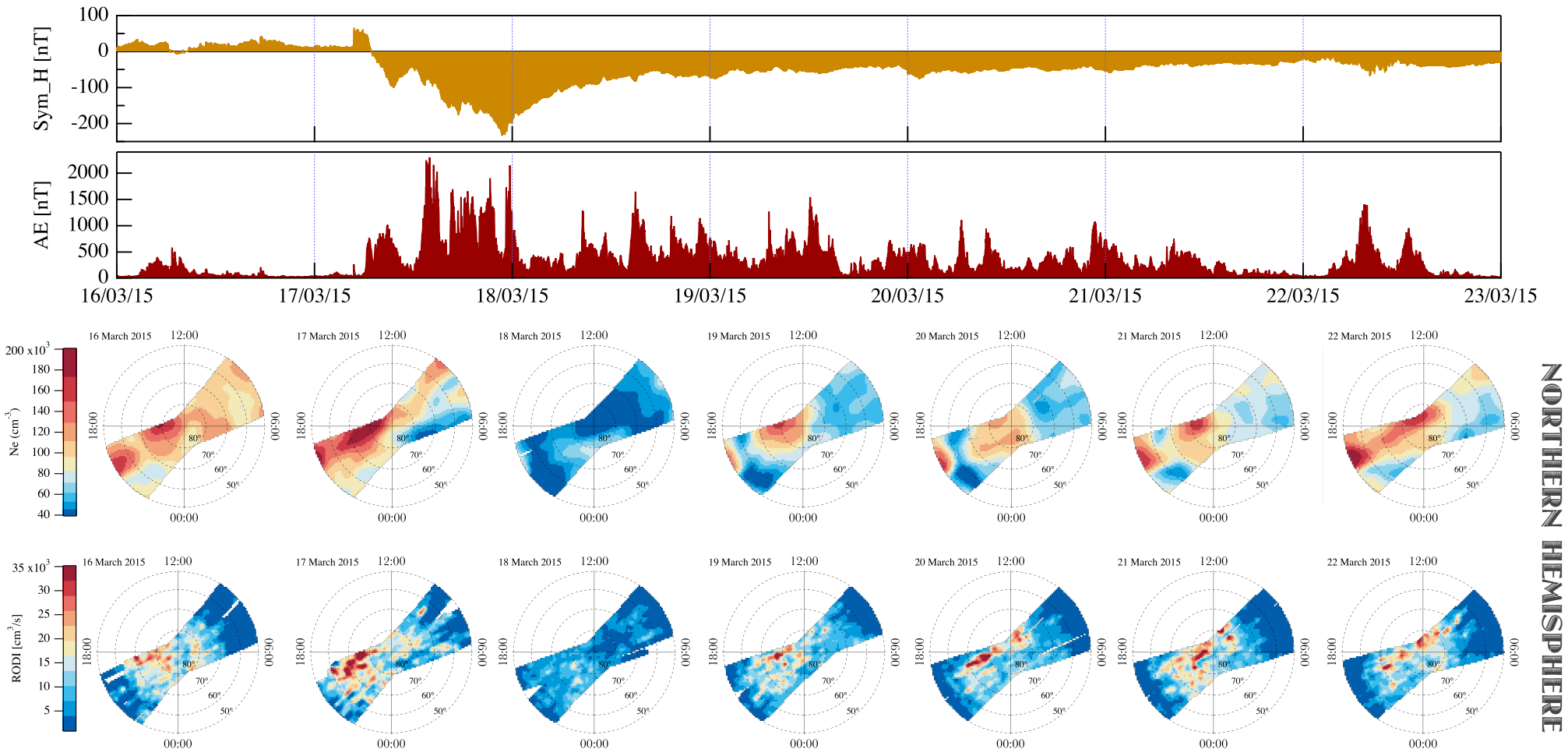


DATASET: Electron Density and RODI (Rate Of change of Density Index) from Swarm constellation



We focus on the analysis of the fluctuations of in-situ electron density provided by the Langmuir probe on board Swarm satellites (data at 1 Hz coming from SwarmA and SwarmB). Data are presented in terms of Quasi-Dipole latitude and MLT coordinates and reported in polar view daily maps.

DATASET: Electron Density and RODI (Rate Of change of Density Index) from Swarm constellation



Using Ne we evaluate the Rate Of change of Density Index (RODI). RODI is commonly used to quantify the structuring of ionospheric plasma and as a proxy of ionospheric plasma irregularities.



Is there a relation between the occurrence of plasma density irregularities as described by the **Rate Of change of electron Density Index** and the **scaling properties of electron density fluctuations** during the development of the geomagnetic storm ?

If the **origin of plasma density irregularities is due to a turbulence process** the plasma density fluctuations are expected to show **scale invariance**.



The study of the scaling properties of electron density fluctuations can be, indeed, important to understand if the **turbulence is the main source of the irregularities and to understand the type of turbulence**.



How to investigate scale-invariance ?

We consider **qth-order structure function** $S_q(\tau)$, which for a signal $N_e(t)$ defined over an interval T is given by

$$S_q(\tau) = \langle |N_e(t + \tau) - N_e(t)|^q \rangle_T$$

when we deal with a scale-invariant signal the $S_q(\tau)$ exhibits a power law behavior as a function of the time separation

$$S_q(\tau) \sim \tau^{\gamma(q)}$$



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To study the **nature of the scaling properties** of electron density fluctuations means to study the **scaling exponents**



How to investigate scale-invariance ?

➔ $S_1(\tau) = \langle |N_e(t + \tau) - N_e(t)| \rangle \approx \tau^{\gamma(1)}$

$\gamma(1)$ is a benchmark of the persistent or antipersistent character of the fluctuations

➔ $S_2(\tau) = \langle |N_e(t + \tau) - N_e(t)|^2 \rangle \approx \tau^{\gamma(2)}$

$\gamma(2)$ describes the spectral properties of the analyzed signal. According to Wiener-Khinchin theorem, the Fourier power spectral density (PSD) exponent β of a signal is related to $\gamma(2)$ according to the following relation:

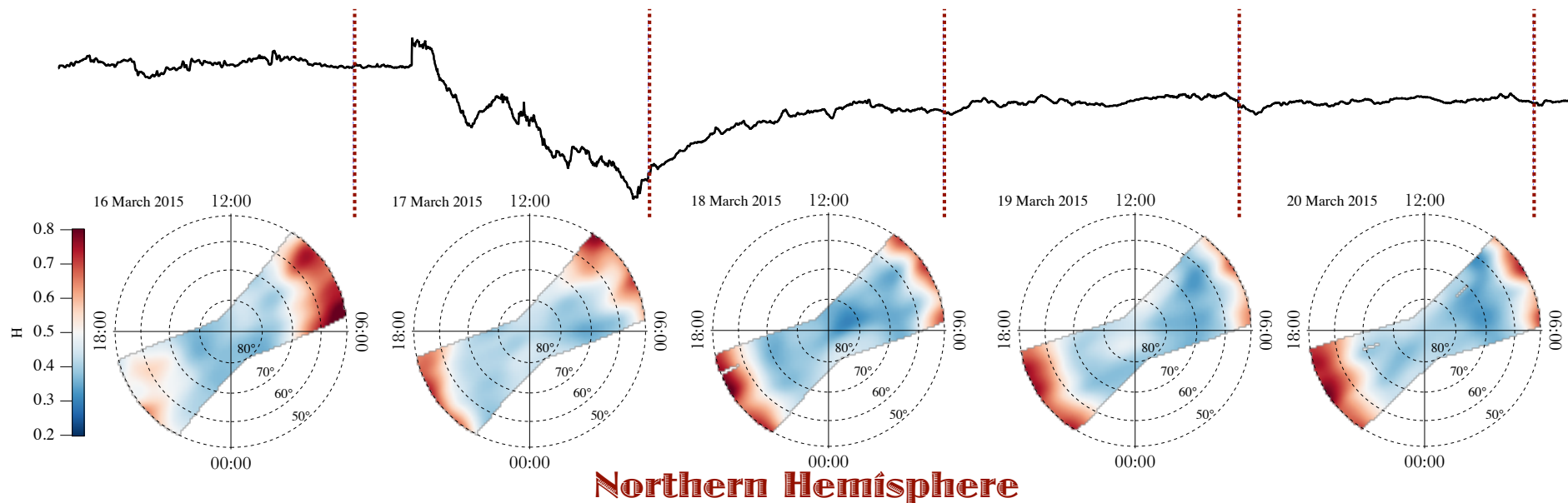
$$PSD(f) \sim f^{-\beta} \rightarrow \beta = \gamma(2) + 1$$



Results: 1st-order Structure Function

$$S_1(\tau) = \langle |N_e(t + \tau) - N_e(t)| \rangle_T \sim \tau^{\gamma(1)}$$

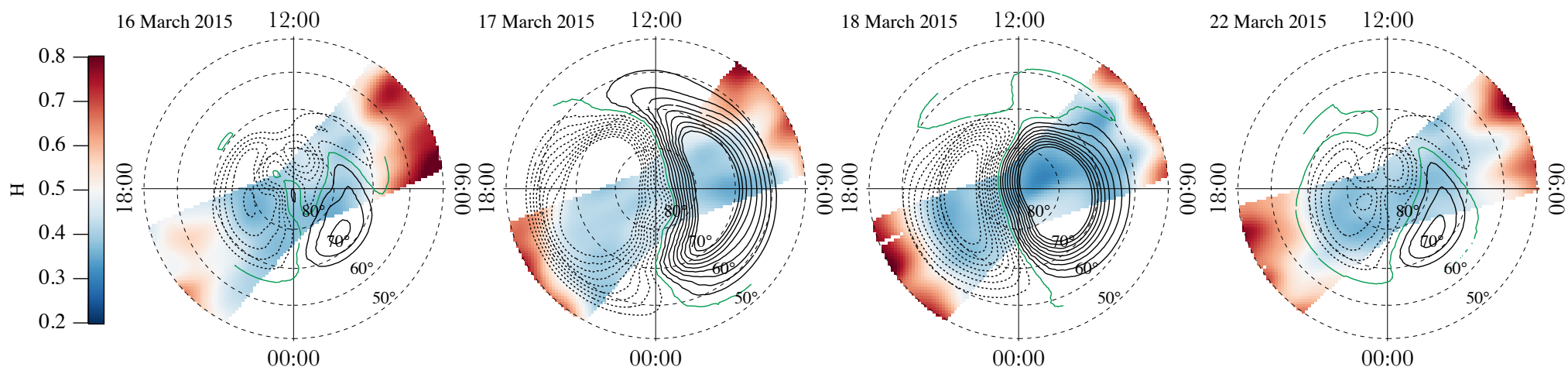
- The **anti persistent character** seems to be a typical features of high-latitudes.
- The **persistent character** appears to be a typical features of mid-latitudes.
- The position of the profile of the local Hurst values with respect to $H = 0.5$ seems to follow the spatial and temporal evolution of the auroral zone during the selected period.





Results: 1st-order Structure Function

Here we have superposed to our results the statistical convection patterns derived from SuperDARN plasma drift measurements and obtained using the “CS10 Statistical Convection Model”.



Northern Hemisphere

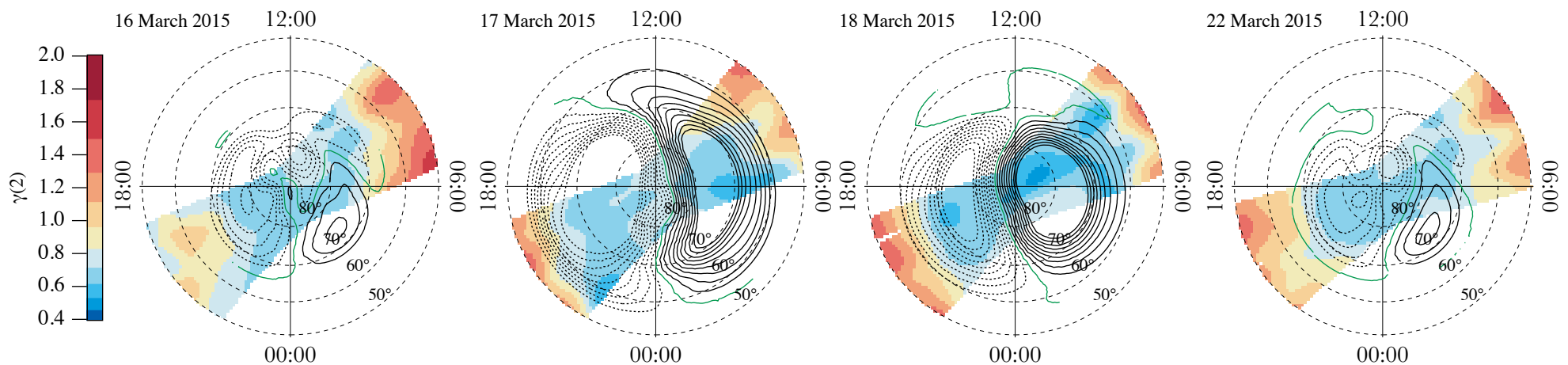
A good agreement is found between the regions characterized by an anti-persistent behavior of the electron density fluctuations and both the position and the extension of the statistical convection patterns obtained in the same period.



Spectral Features: Spectral Density Exponent

$$S_2(\tau) = \langle |N_e(t + \tau) - N_e(t)|^2 \rangle \approx \tau^{\gamma(2)}$$

Also in this case, we have superposed to our results the statistical convection patterns derived from SuperDARN plasma drift measurements and obtained using the “CS10 Statistical Convection Model”.



Northern Hemisphere

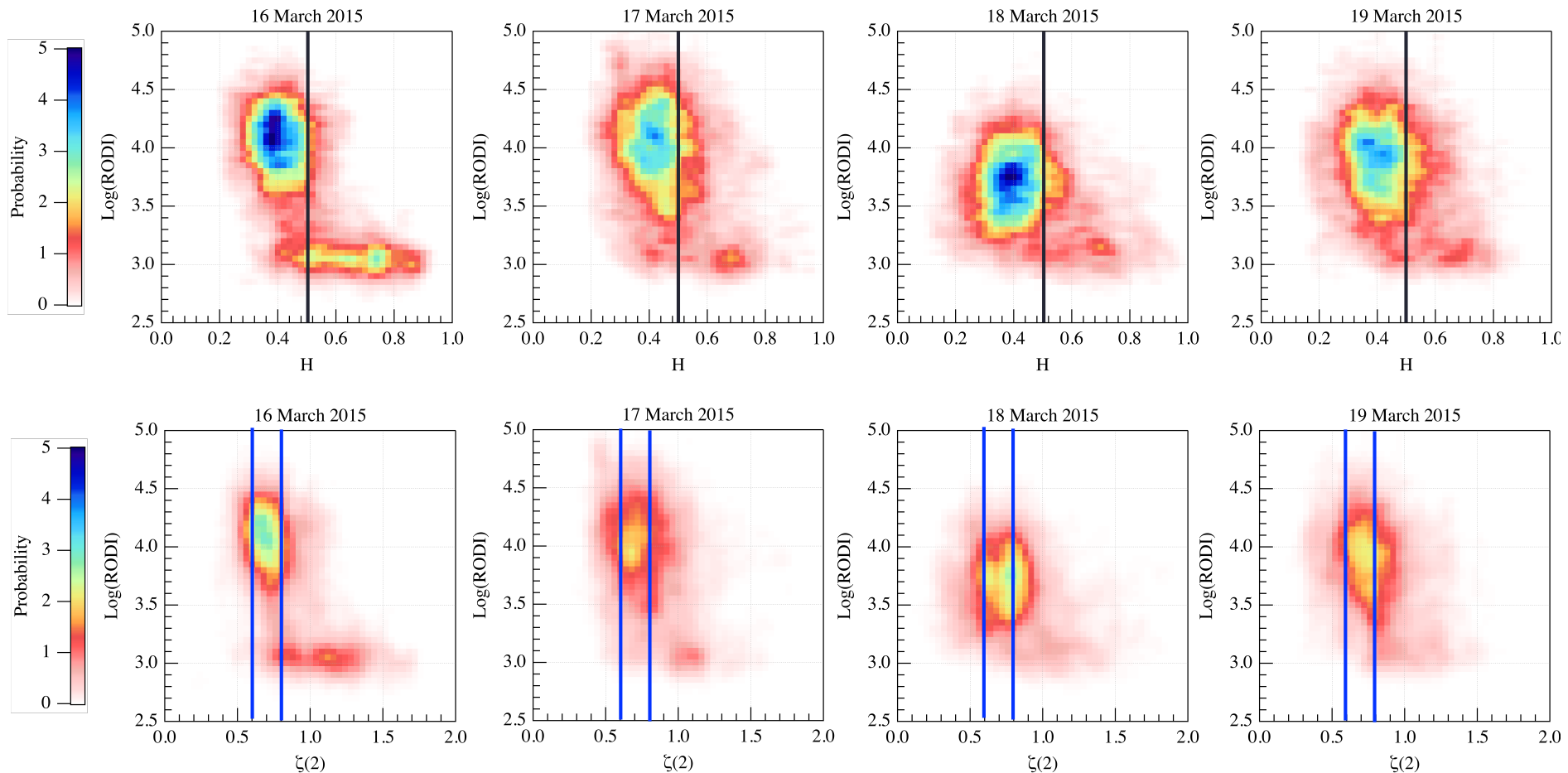
Different scaling (turbulence) regimes and processes may characterize the regions where the convection patterns are localized. The values in the range 0.6 to 0.8 support the idea of a fluid and/or MHD turbulence as a source of these perturbations.



Results:

RODI vs Structure Function Exponents of Electron Density

To investigate the possible relationship between RODI parameter and the scaling features of the Ne fluctuations we evaluate the **joint probability distributions** between RODI and first and second-order scaling exponents.



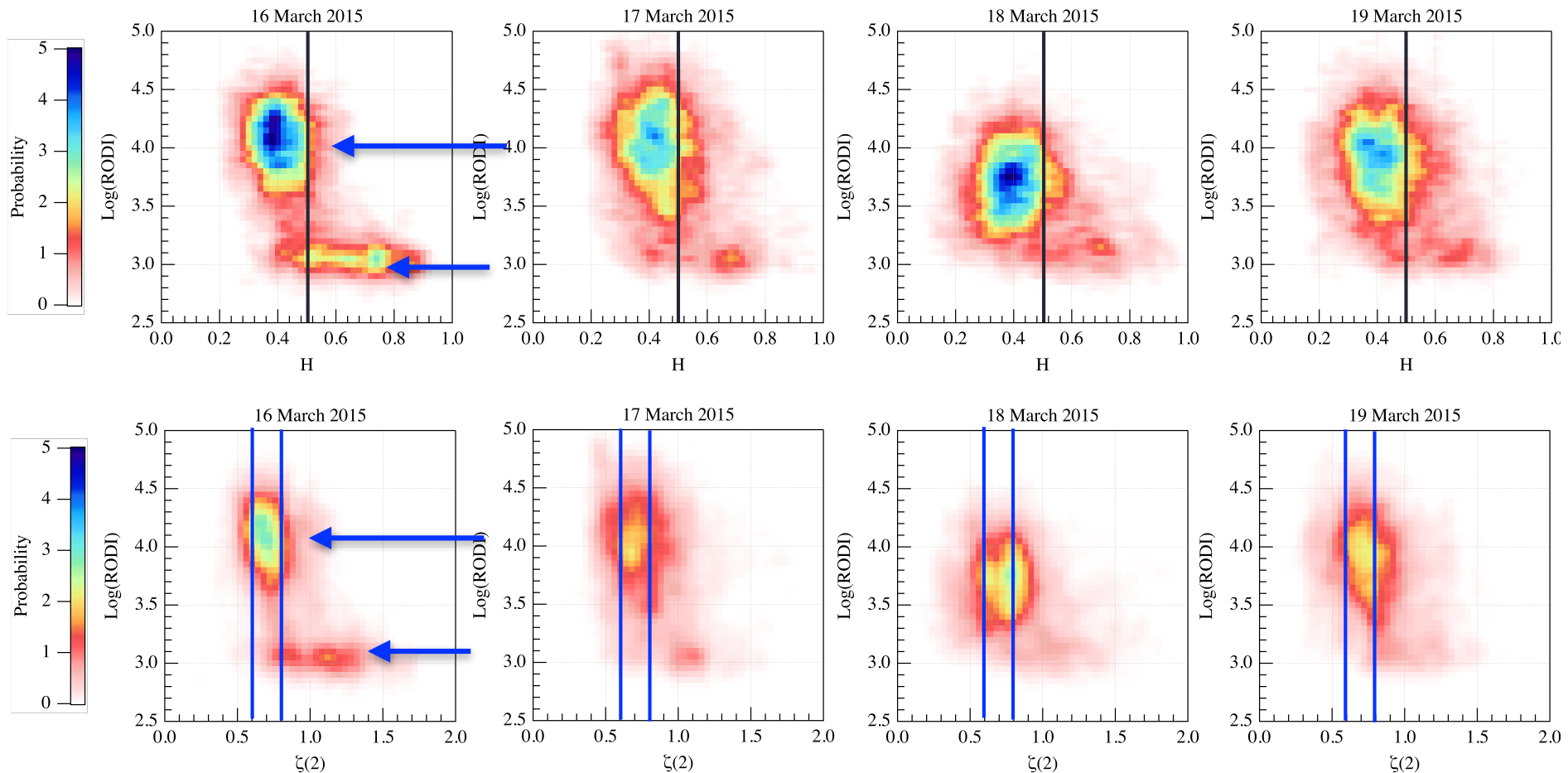
Northern Hemisphere



Results:

RODI vs Structure Function Exponents of Electron Density

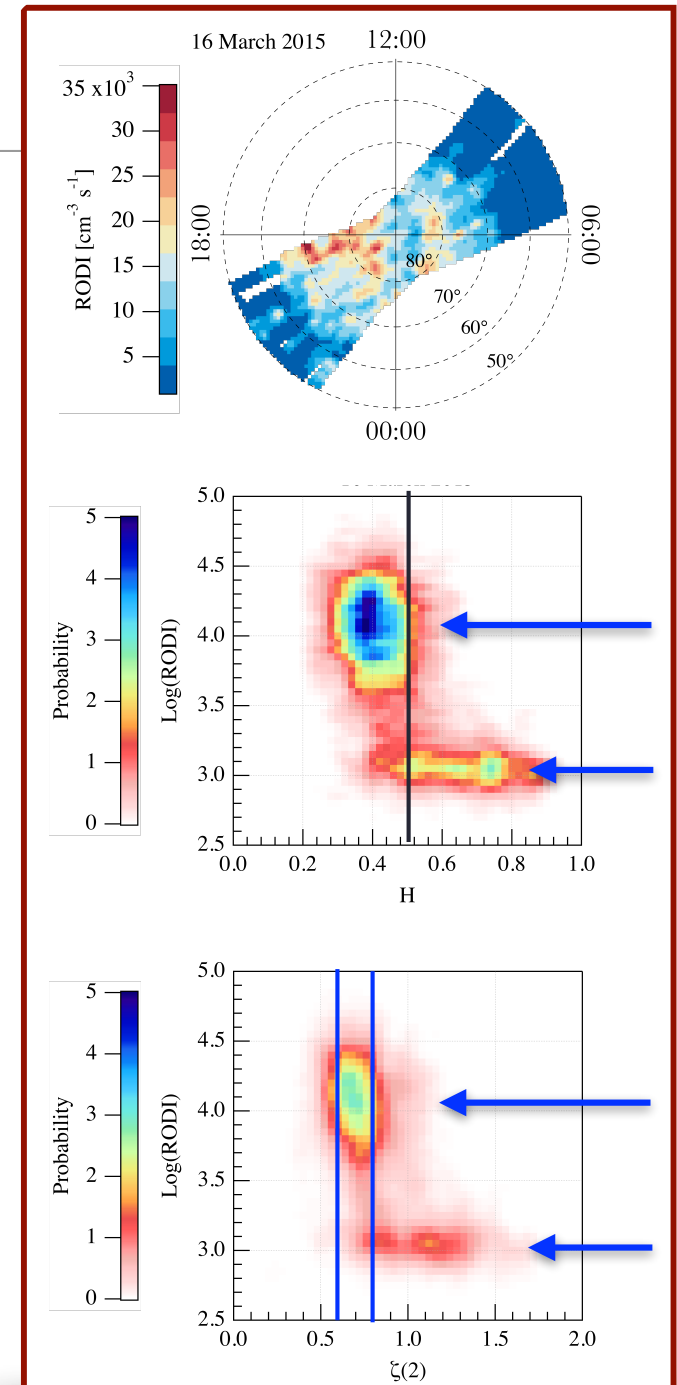
The scaling features of electron density fluctuations permit us to find **two families** of plasma irregularities which seem to be associated with different physical properties.



Northern Hemisphere

Summarizing ...

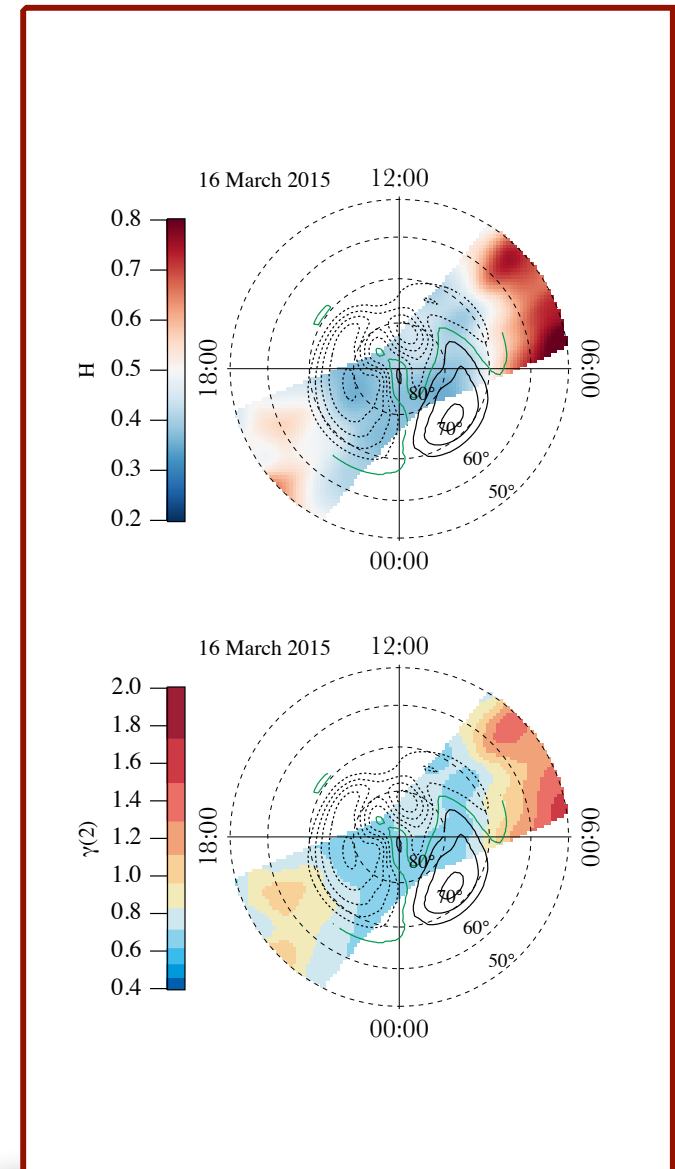
- The plasma density variations may be due to different processes not necessarily linked to turbulence phenomena.
- The RODI parameter cannot be generically considered as a proxy of the occurrence of ionospheric turbulence
- The scaling features of electron density fluctuations permitted us to find two families of plasma irregularities which seem to be associated with different physical properties.
- Only **plasma density variations associated with high values of RODI are characterized by scaling properties which support the idea of a fluid and/or MHD turbulence as a source of these perturbations.**





Summarizing ...

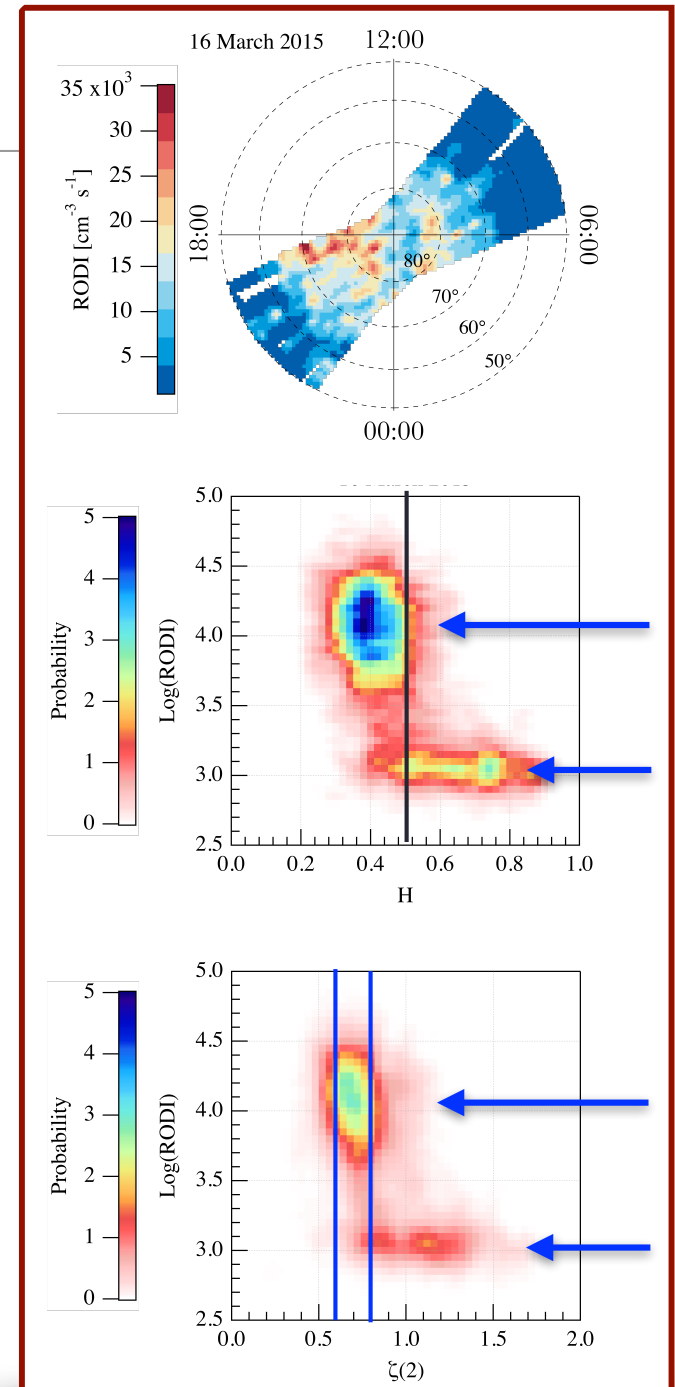
- The scaling properties of the electron density fluctuations slightly change during the development of the geomagnetic storm, depend on the MLT and latitude values.
- A good agreement is found between the regions characterized by a **antipersistent behavior** of the electron density fluctuations and both the position and the extension of the statistical **convection patterns** obtained in the same period using SuperDARN data.
- The regions characterized by an anti persistent behavior of the fluctuations are also characterized by values of around 0.6-0.8





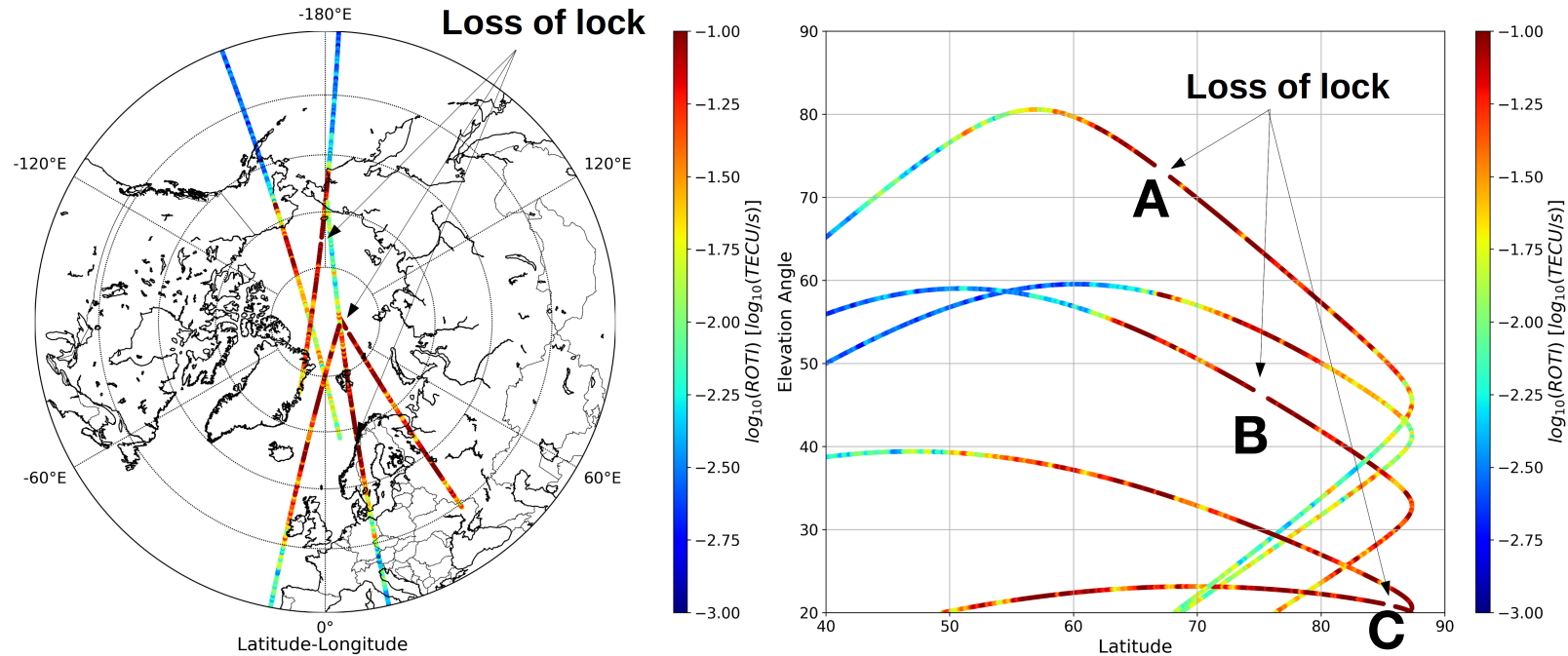
Summarizing ...

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Work is going on ...

Scatter ROTI, 17 March 2015, North Pole, Swarm A, PRN=25

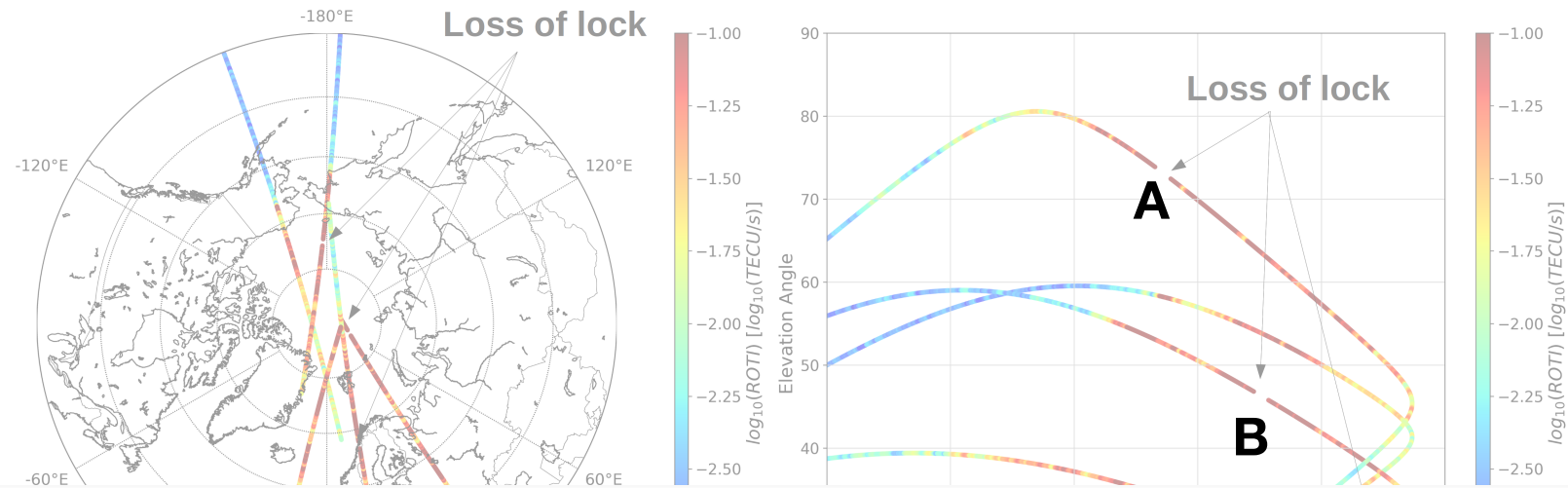


	Case A	Case B	Case C
1st scaling exponent	0.39 ± 0.04	0.28 ± 0.03	0.49 ± 0.05
2nd scaling exponent	0.67 ± 0.07	0.43 ± 0.04	0.82 ± 0.08

Electron density variations associated with loss of lock are characterized by scaling properties which support the idea of a fluid and/or MHD turbulence as a source of these phenomena.

Work is going on ...

Scatter ROTI, 17 March 2015, North Pole, Swarm A, PRN=25



These results pave the way for the elaboration of a future product capable of providing information on the possible development of those turbulent phenomena which can have repercussions in the space weather framework

Electron density variations associated with loss of lock are characterized by scaling properties which support the idea of a fluid and/or MHD turbulence as a source of these phenomena.

Thank you!

